



# Simulation and analysis of HWP and TES non-idealities for CMB space missions

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*CMB-CAL @ Bicocca, Milan*

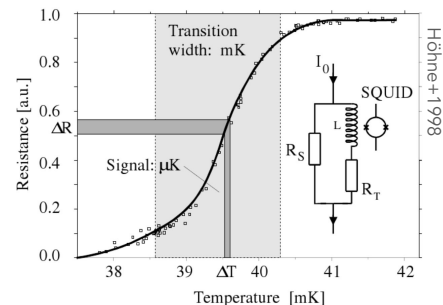
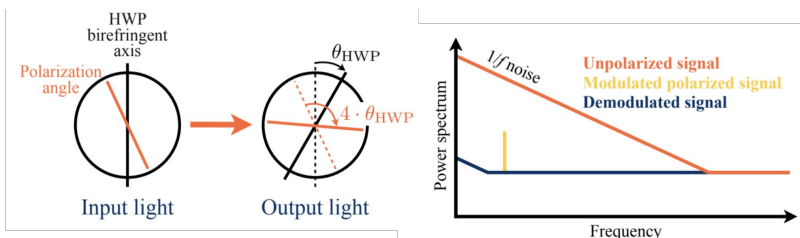
# Outline of this talk

- Introduction
- Analytical model of the systematics
- Simulations applied to LiteBIRD
- Discussion and conclusions

# HWP and TES in a nutshell:

- A RHWP moves CMB polarization signal above  $1/f$  noise, reducing systematic effects
- Fast HWP rotation allows quick single-detector measurement of  $I$ ,  $Q$ ,  $U$   $\rightarrow$  minimizes need for corrections due to detector mismatches

- A TES is a cryogenic super sensitive thermistor
- It measures current at a fixed voltage; current changes reflect optical power through the resistance shift in a superconducting film.



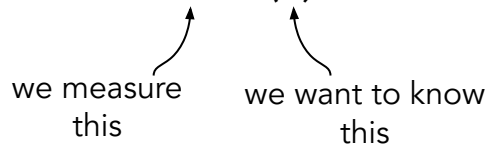
# Parasitic signals from HWP systematics

- Parasitic signals at harmonics of HWP rotation frequency ( $f_{\text{HWP}}$ ) can be caused by HWP imperfections.
- HWPSS correction or calibration methods studied in previous experiments (e.g., NIKA, ABS, POLARBEAR).
- HWPSS are usually modeled as: 
$$A(\chi) = \sum_{n=1}^{\infty} (A_n + a_n^{\text{opt}} I) \cos(n\chi + \phi_n)$$
- Focus on signal at  $2f_{\text{HWP}}$ :
  - Arises from differential transmission and emission along the HWP's ordinary and extraordinary axes.
  - Can bias signal reconstruction when coupled to detector non-linearity

# TES bolometers non-linearity

Tommaso's talk for more details!

- Each TES detector has some response function  $S$  such that  $I = S(P)$



- Commonly  $S$  is just assumed as a *linear* gain, which is fine in a small-signal approximation (CMB is small enough)
- If the signal is not-so-small (Galaxy/planets crossing, CMB dipole, HWPSS)  $\rightarrow S$  enters a non-linear regime

# TES bolometers non-linearity

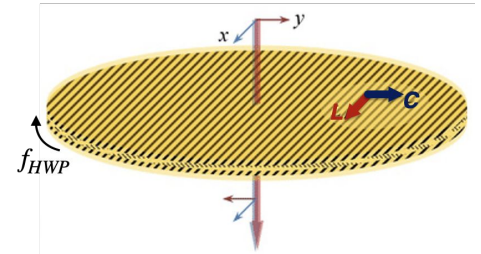
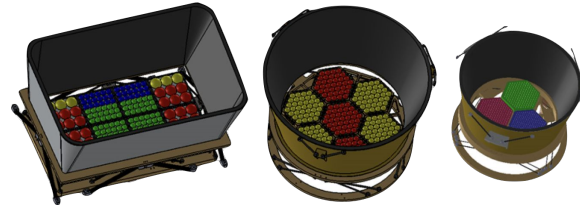
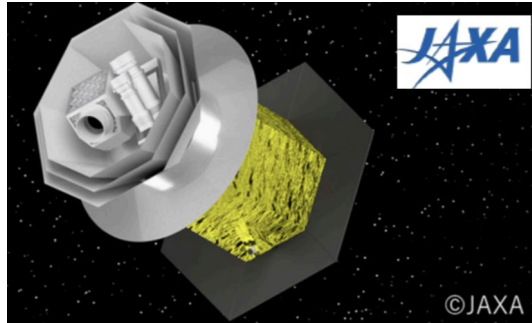
- Each TES detector has some response function  $S$  such that  $I = S(P)$

we measure this  $\nearrow$   $I = S(P)$   $\nwarrow$  we want to know this

- Commonly  $S$  is just assumed as a *linear* gain, which is fine in a small-signal approximation (CMB is small enough)
- If the signal is not-so-small (Galaxy/planets crossing, CMB dipole, HWPSS)  $\rightarrow$   $S$  enters a non-linear regime

*What happens when the HWPSS induces TES non-linearity?  
Which level of non-linearity can we tolerate?*

# Application: the LiteBIRD experiment



## Objectives of this work:

- Evaluate the impact of HWP non-idealities in terms of HWPSS for LiteBIRD baseline configuration
- Optimize MHFT HWPs design to minimize differential emission/transmission
- Determine desired residual non-linearity knowledge to ensure accurate estimation of [\(LiteBIRD collaboration+23\)](#)

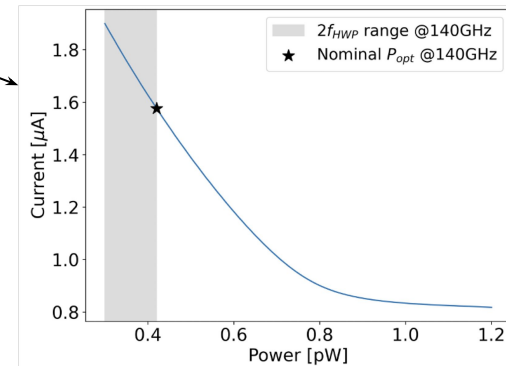
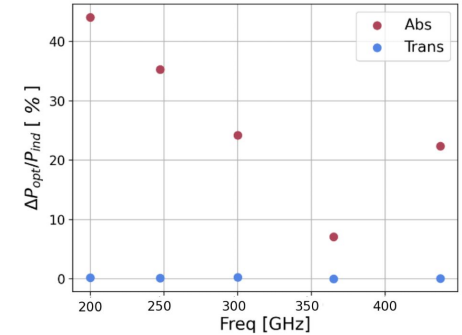
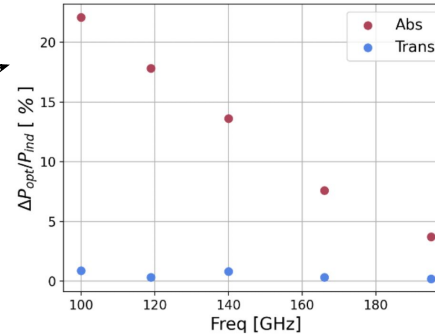
# Application: the LiteBIRD experiment

Pipeline described in [Micheli+2024](#)

- We build our timeline including HWPSS from differential emission at  $2f_{\text{HWP}}$
- We model TES non-linearity as a second order gain correction:  

$$d_{NL}(t) = [1 + g_1 d(t)]d(t)$$
- TES non-linearity up-modulates the  $2f_{\text{HWP}}$  signal resulting in a  $1 \rightarrow 4f_{\text{HWP}}$  leakage

Relative optical power variation in MHFT channels with 20K HWP (baseline config)



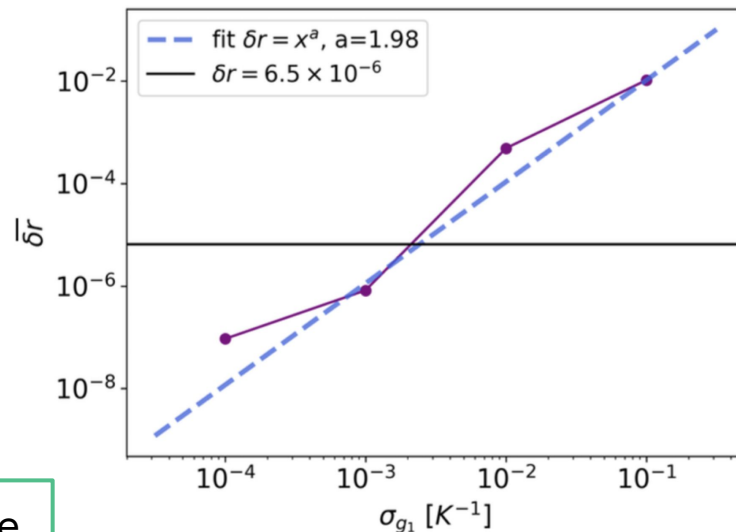


# Non-linearity knowledge requirement for MHFT

- We chose to fix the HWPSS amplitude (as in the baseline configuration) to find the acceptable non-linearity residual
- We apply a non-linear gain for a pair of orthogonal detectors in each MHFT channel with:

$$d_{NL}^i(t) = d^i(t)[1 + \mathcal{N}(0, \sigma_{g_1})]$$

We show that it is crucial to properly account for the coupling between HWP differential emission and TES non-linearity in order not to bias the estimation of  $r$



# Conclusions

- Modelization and injection of HWPSS and non-linearity in the time streams → **done**
- Propagation up to the estimation of the bias on  $r$  to derive non-linearity knowledge requirement → **done for baseline configuration**
- Modelling of TES non-linearity for LiteBIRD → done in [de Haan+24](#)
- Optimization of HWP design to minimize the impact of HWPSS from differential optical properties → **ongoing**

*Thank you!*

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