

Bandpass Calibration with the Frequency-selectable **L**aser **S**ource (FLS)

Shreya Sutariya
On behalf of the Simons Observatory Collaboration

Bandpass Calibration with the FLS

Motivation

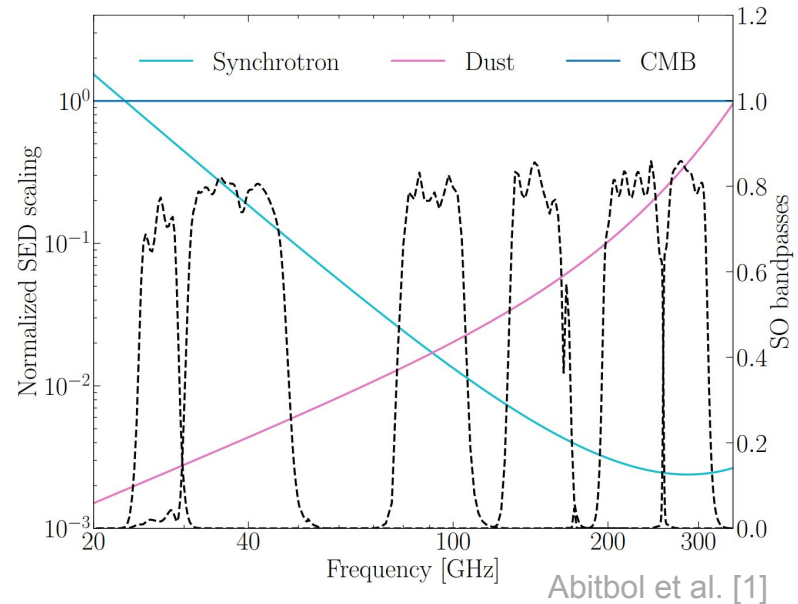
Frequency-selectable Laser Source (FLS) Instrument Design

FLS Characterization

Measurement Procedure and Results

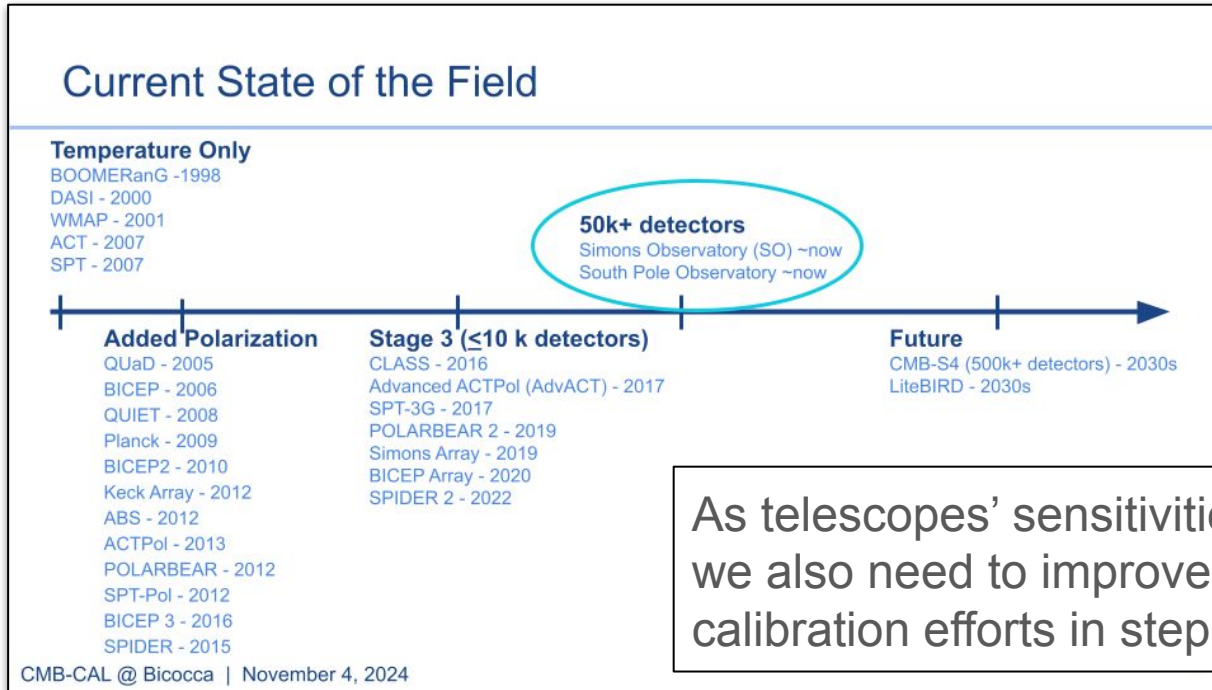
Bandpass Calibration and Component Separation

- Astrophysical components like synchrotron and dust also emit in the CMB frequencies.
- Each component has its own frequency dependence, its Spectral Energy Distribution.
- Component separation requires knowledge of the detectors' bandpasses to successfully isolate the different signals.



1. Maximilian H. Abitbol et al., “The Simons Observatory: gain, bandpass and polarization-angle calibration requirements for B-mode searches,” JCAP05 (2021) 032

Bandpass Calibration for next-generation CMB experiments



Credit: S. Simon

Bandpass Calibration for next-gen CMB experiments

- Currently, Fourier Transform Spectrometers (FTSes) achieve a few percent-level accuracy of the bandpasses [T. Alford et al., in prep]
- Science drivers like inflation and cluster science for current and future experiments put more stringent requirements on bandpasses.

Outline

Motivation

Frequency-selectable Laser Source (FLS) Instrument Design

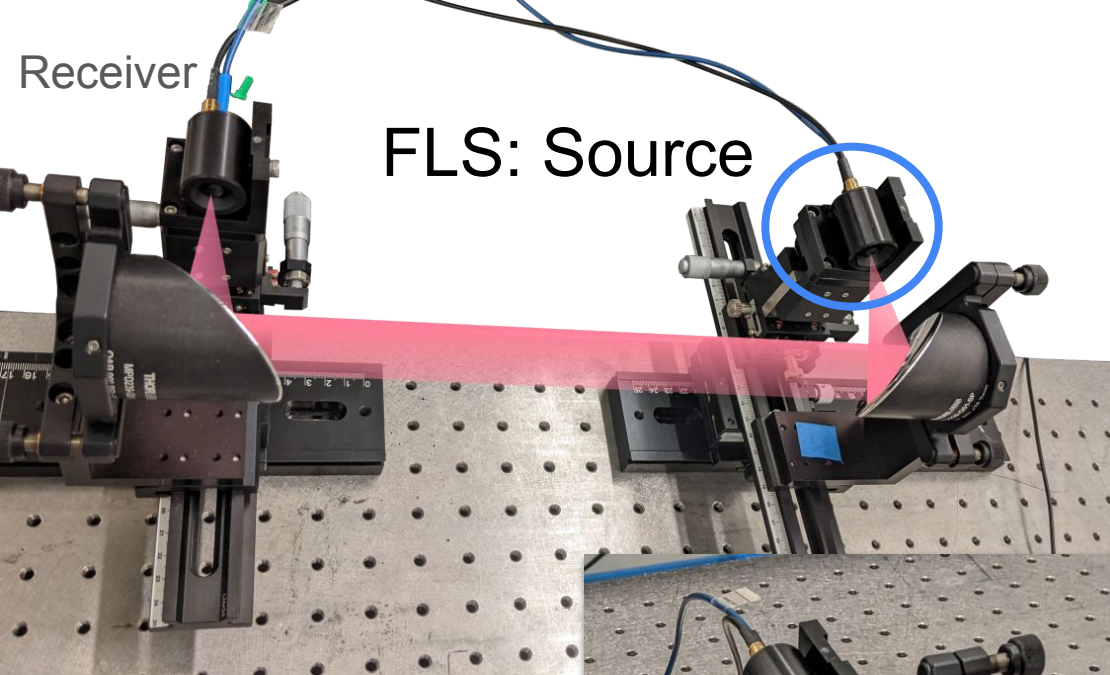
FLS Characterization

Measurement Procedure and Results

- The FLS is a calibration instrument to measure bandpasses and out-of-band leakage.
- FLS and FTS measurements show agreement, with the FLS being able to measure edges more sensitively.

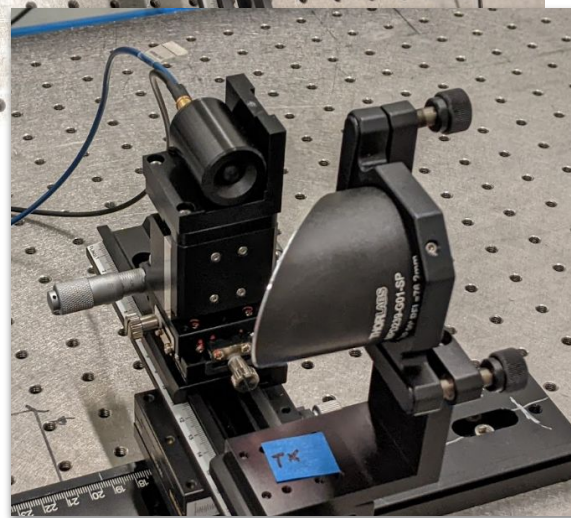


1. Source
2. Attenuators
3. Coupling Optics



Commercially available GHz/THz laser that emits from 20 - 1200 GHz with fine resolution.

- The laser system consists of photomixer source and receiver.
- Emits $\sim 65 \mu\text{W}$ at 100 GHz.

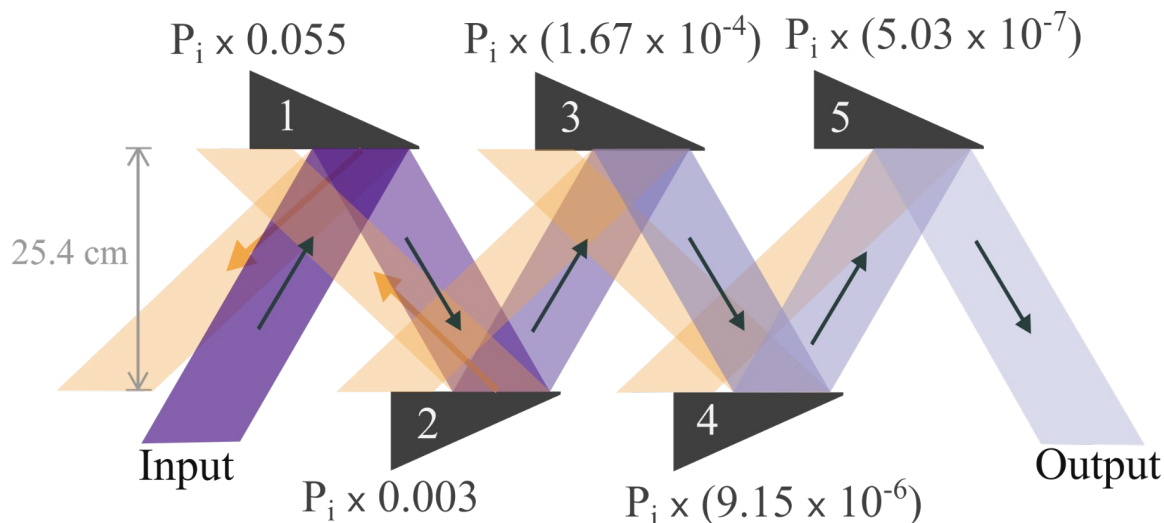


1. Source
2. **Attenuators**
3. Coupling Optics

Prism Attenuators + Power Variability

- Incident light loses ~94% of its power after one reflection
- Using up to five prisms can significantly attenuate the laser power.

Each prism is replaceable by a reflective plate, allowing for power tunability.



Coupling Optics

1. Source
2. Attenuators
3. Coupling Optics

Output/Receiver Module

Source Module

Receiver for alignment

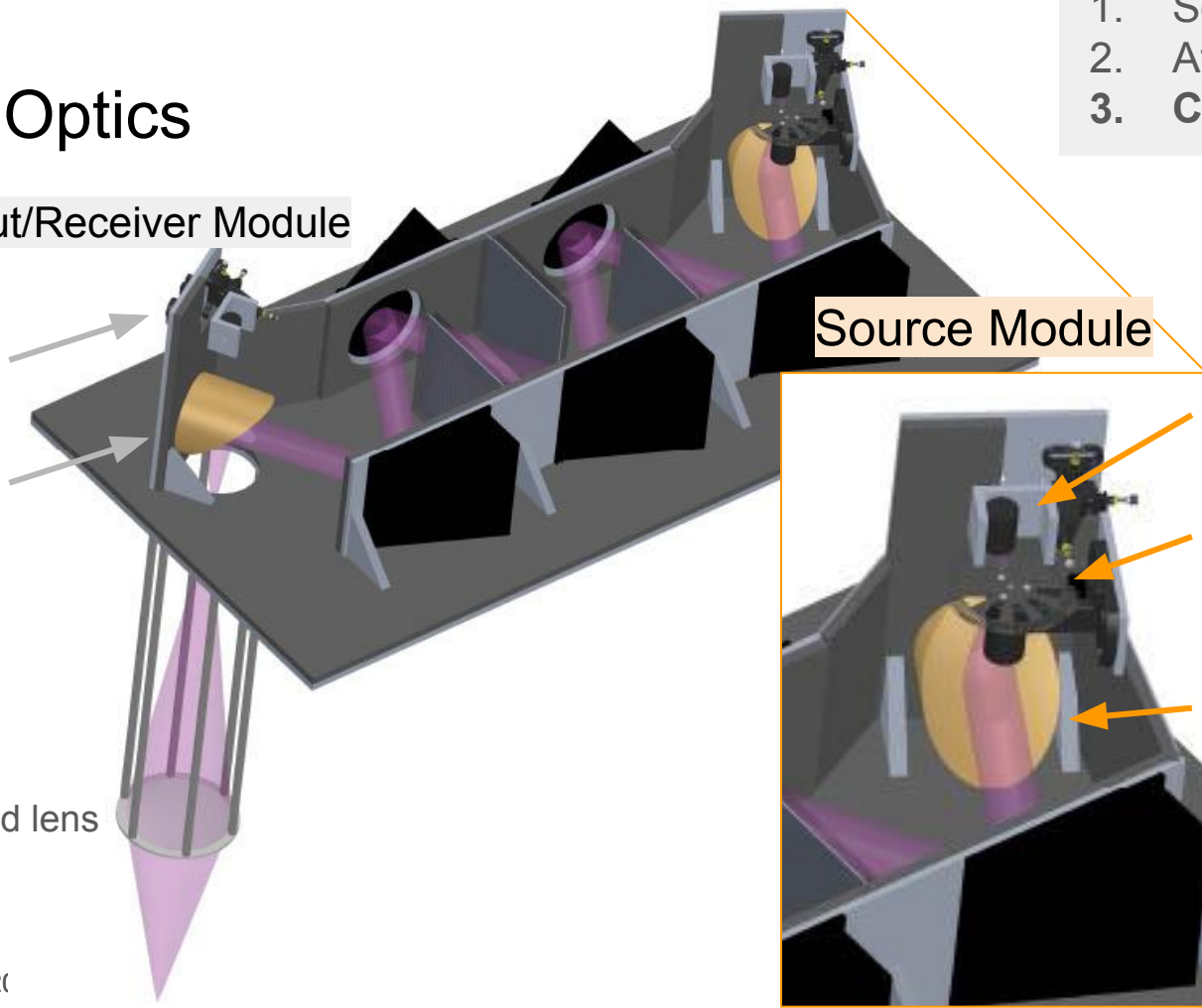
Off-axis Parabolic Mirror

Double sided lens

Source

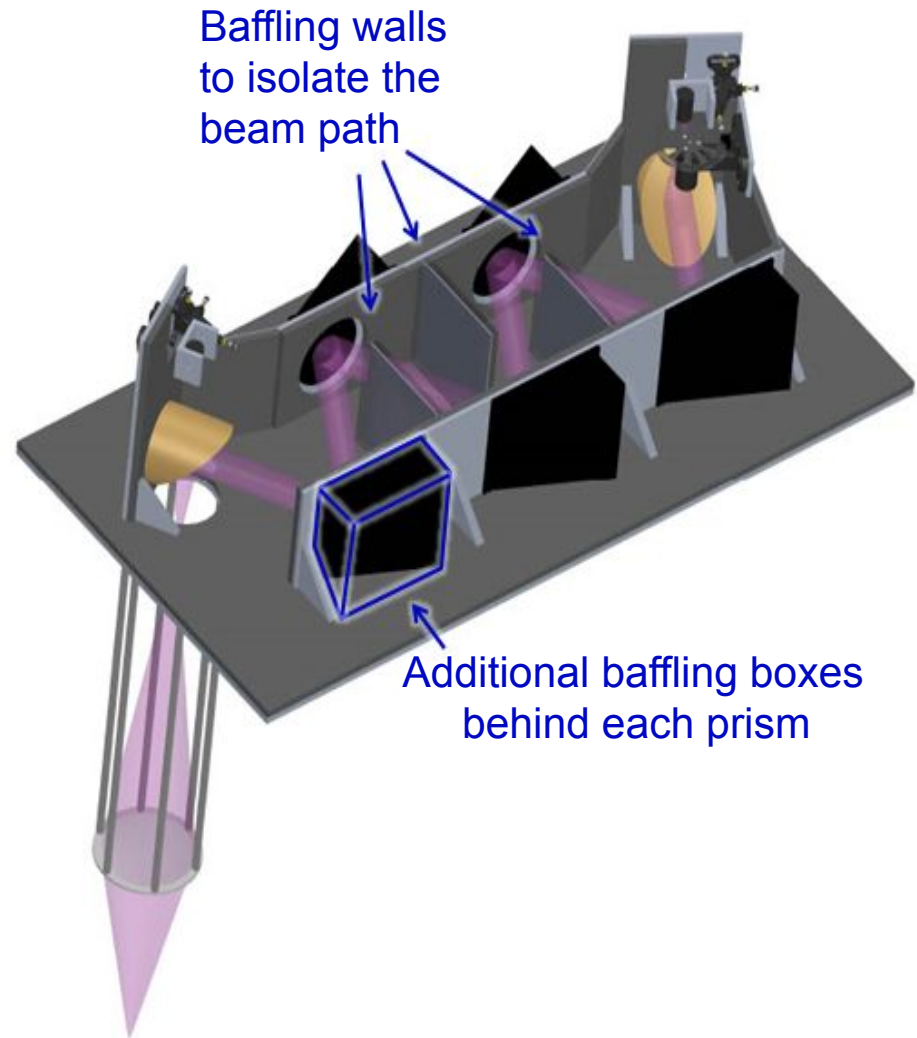
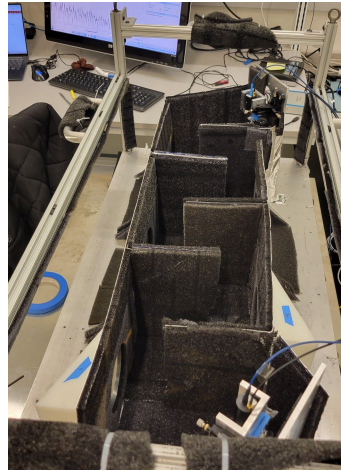
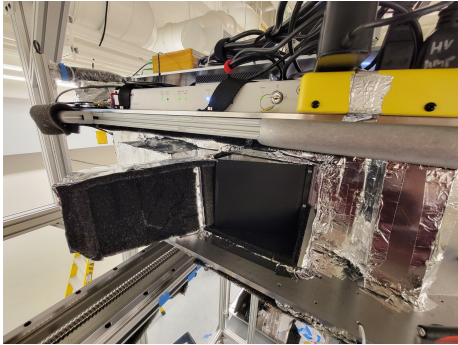
Chopper

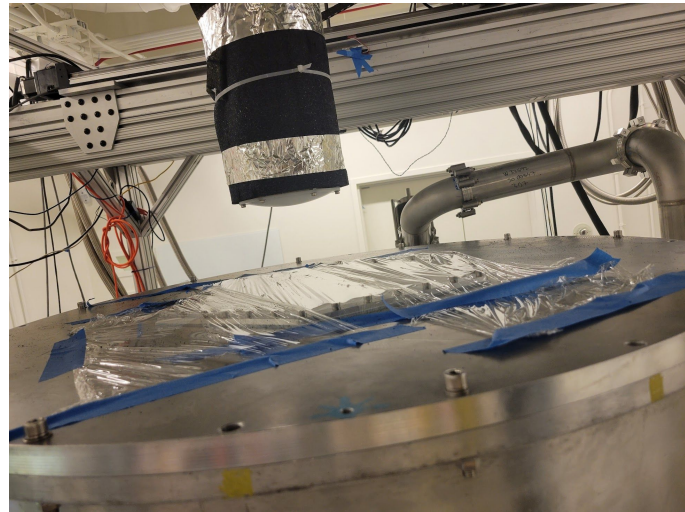
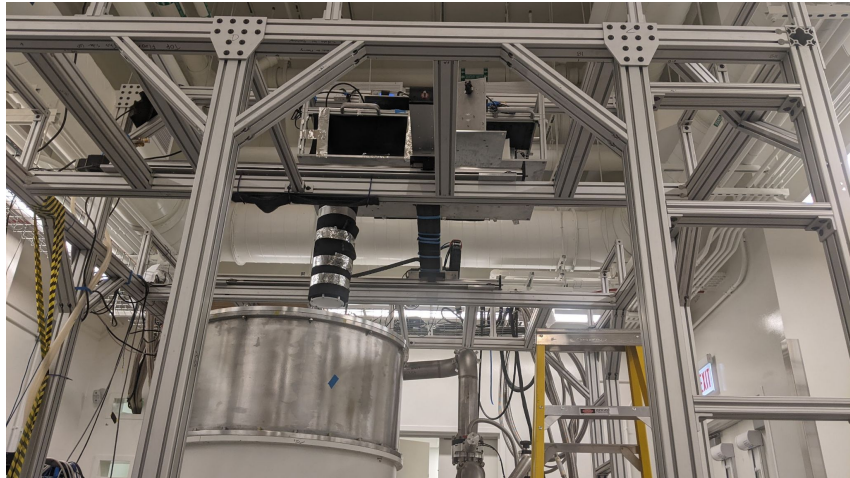
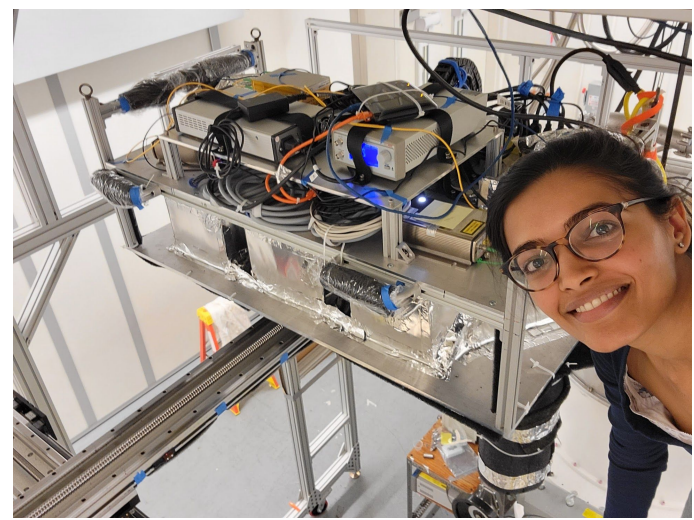
Off-axis Parabolic Mirror



Absorptive Baffling

As a result of attenuating away so much power, the system becomes more vulnerable to unwanted scattering and other systematics.





CMB-CAL Milano Bicocca 2024

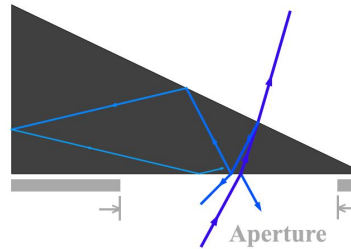
Outline

Motivation

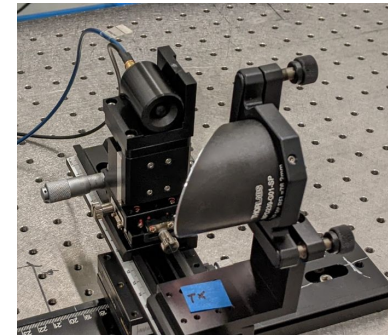
Frequency-selectable Laser Source (FLS) Instrument Design

FLS Characterization

Measurement Procedure and Results

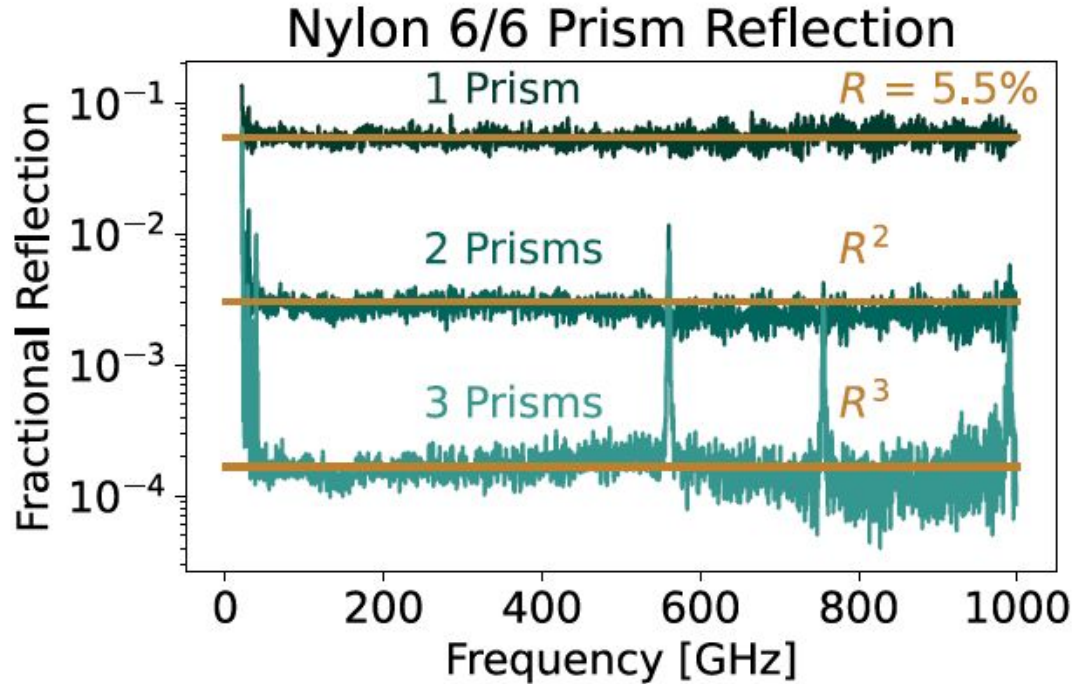


Prism Reflectance
as a function of
frequency



Source Frequency
and Power
Characterization

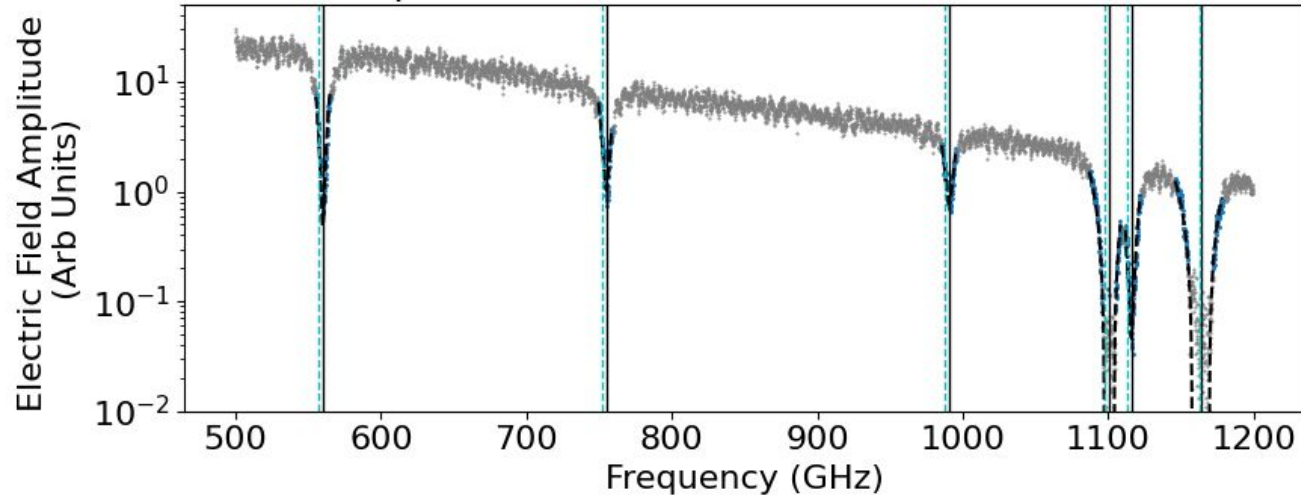
Prism Reflectance



The prisms provide a neutral attenuation across a wide frequency range.

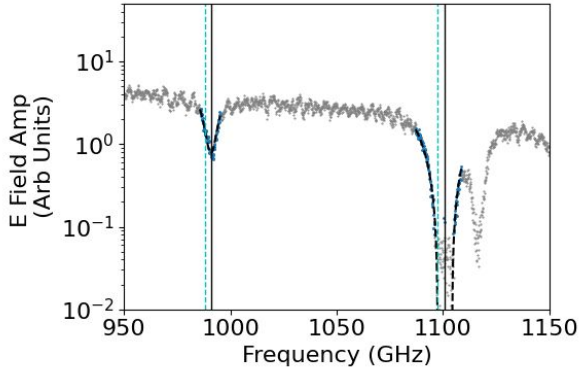
Frequency Calibration

Atmospheric Transmission measured with Laser

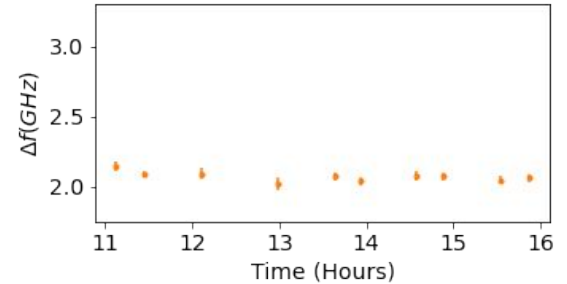
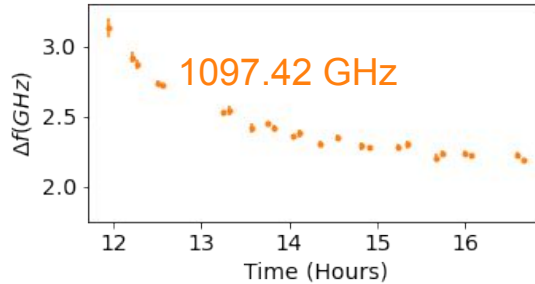
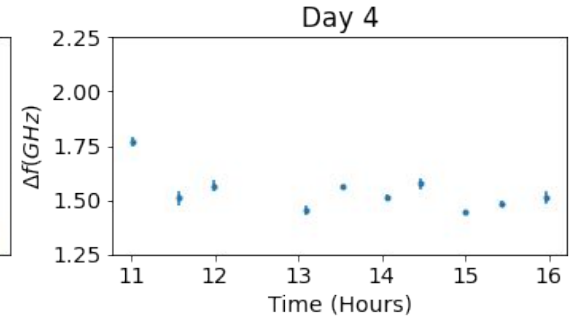
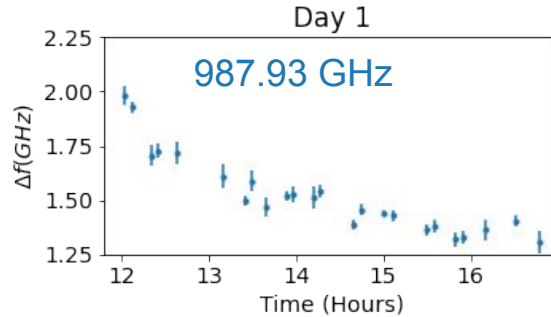


We can measure the frequency accuracy by measuring the water vapor absorption lines and comparing them to known values.

Frequency Settling Time

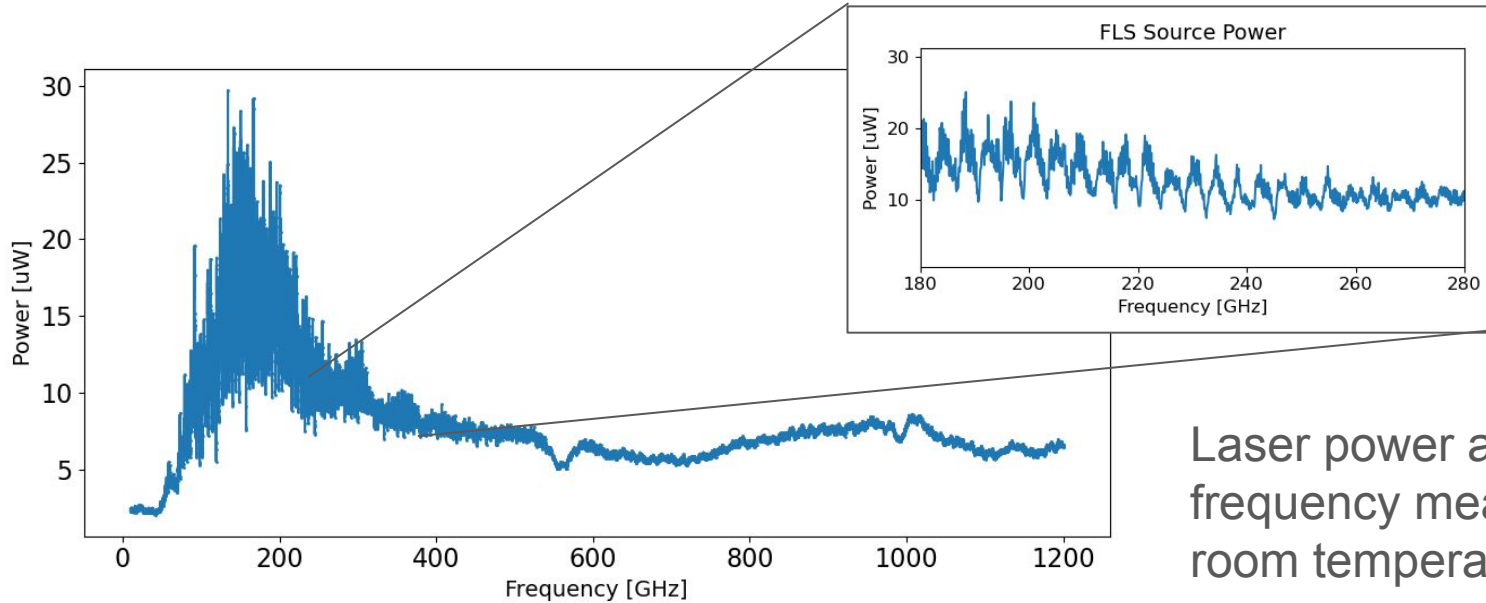


Frequencies settle after ~36 hours and retain a constant offset that can be corrected.



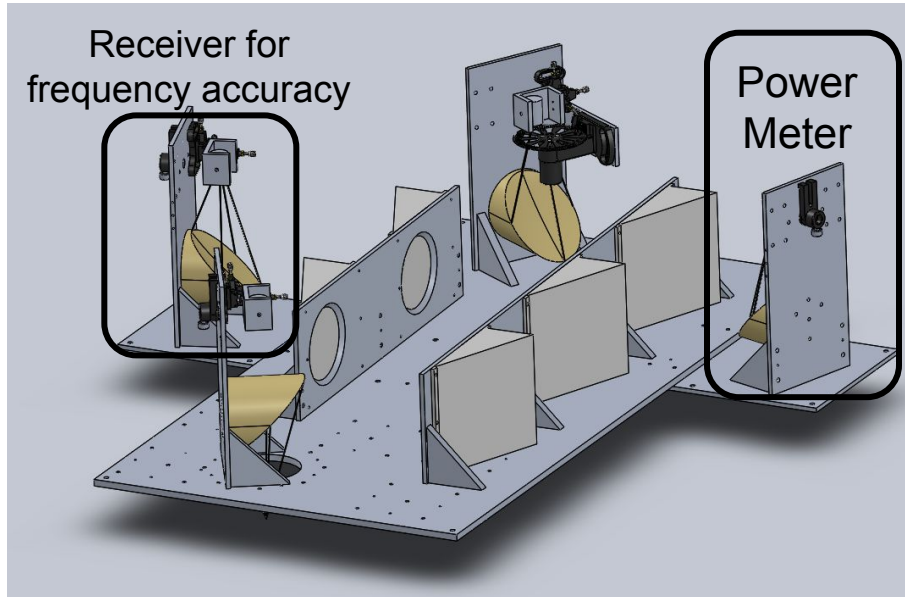
Credit: Lauren Saunders

Power Calibration Measurement



Laser power as a function of frequency measured with a room temperature bolometer.

Design for in-situ calibration modules



In-situ modules
incorporated into next
version of FLS!
Upgrades in progress by
L. Saunders and S.
Simon

Outline

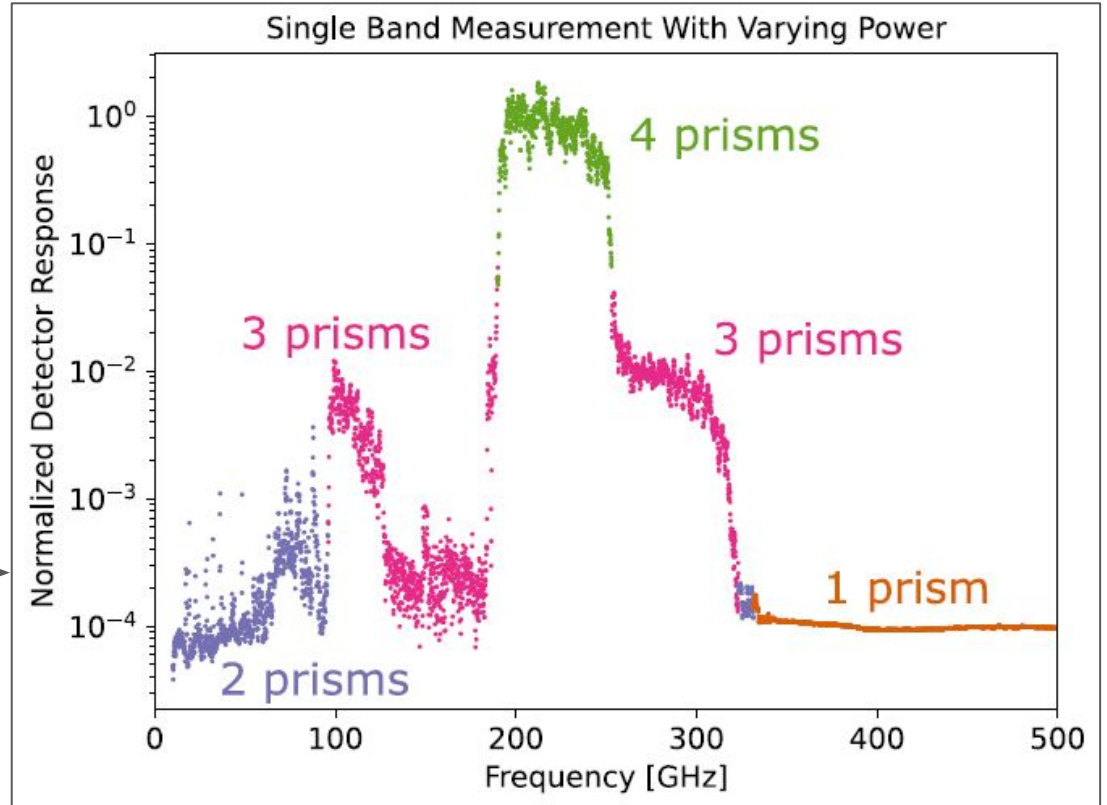
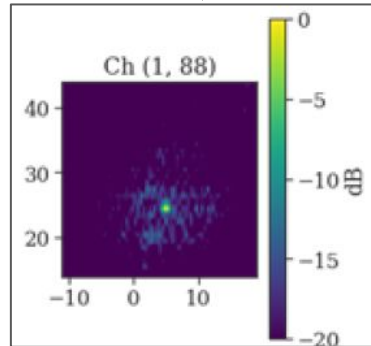
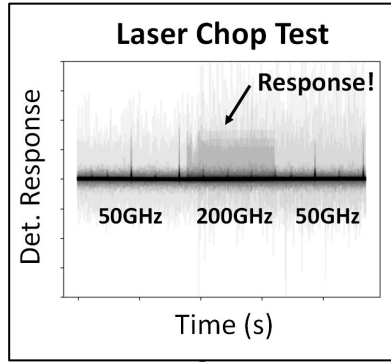
Motivation

Frequency-selectable Laser Source (FLS) Instrument Design

FLS Characterization

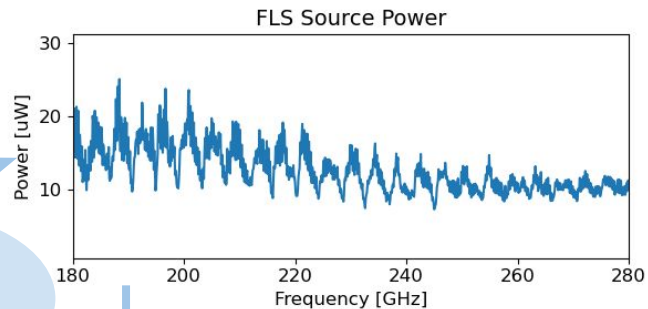
Measurement Procedure and Results

Beam Maps and Bandpass Measurements



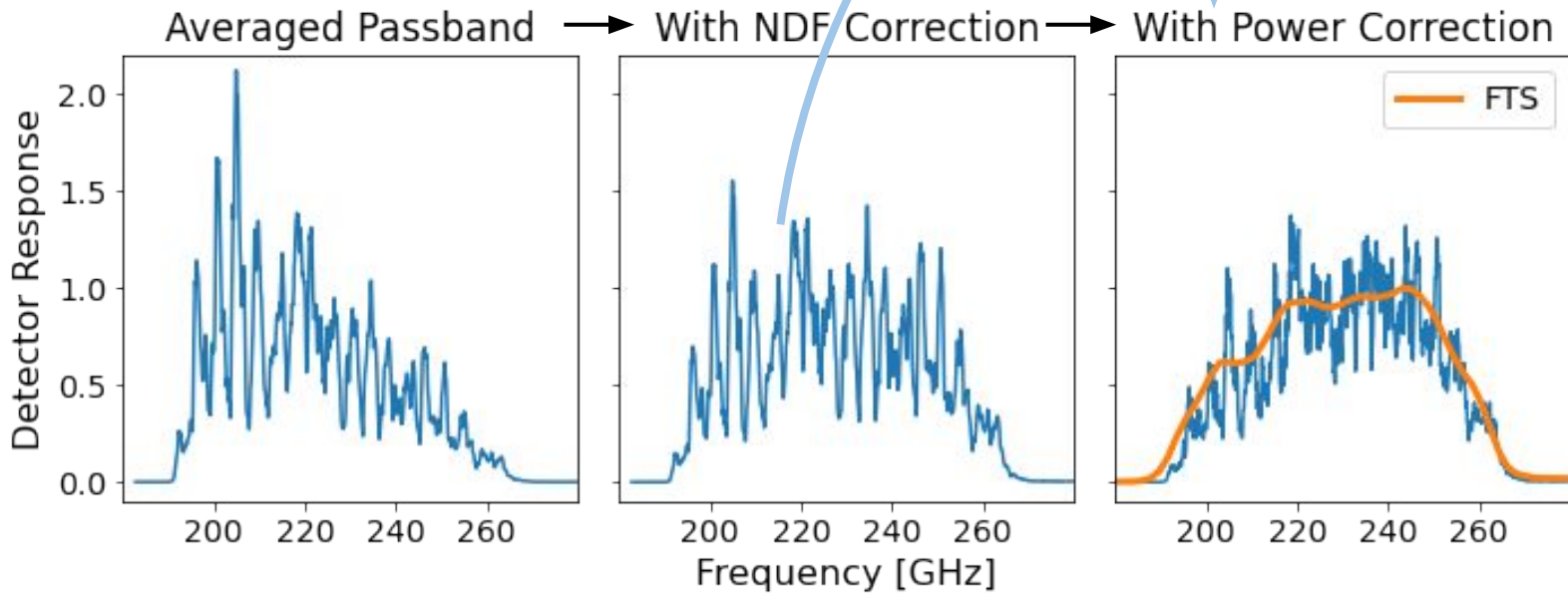
Post-measurement Corrections

- non-Neutral Density Filter (NDF) correction
- Power correction



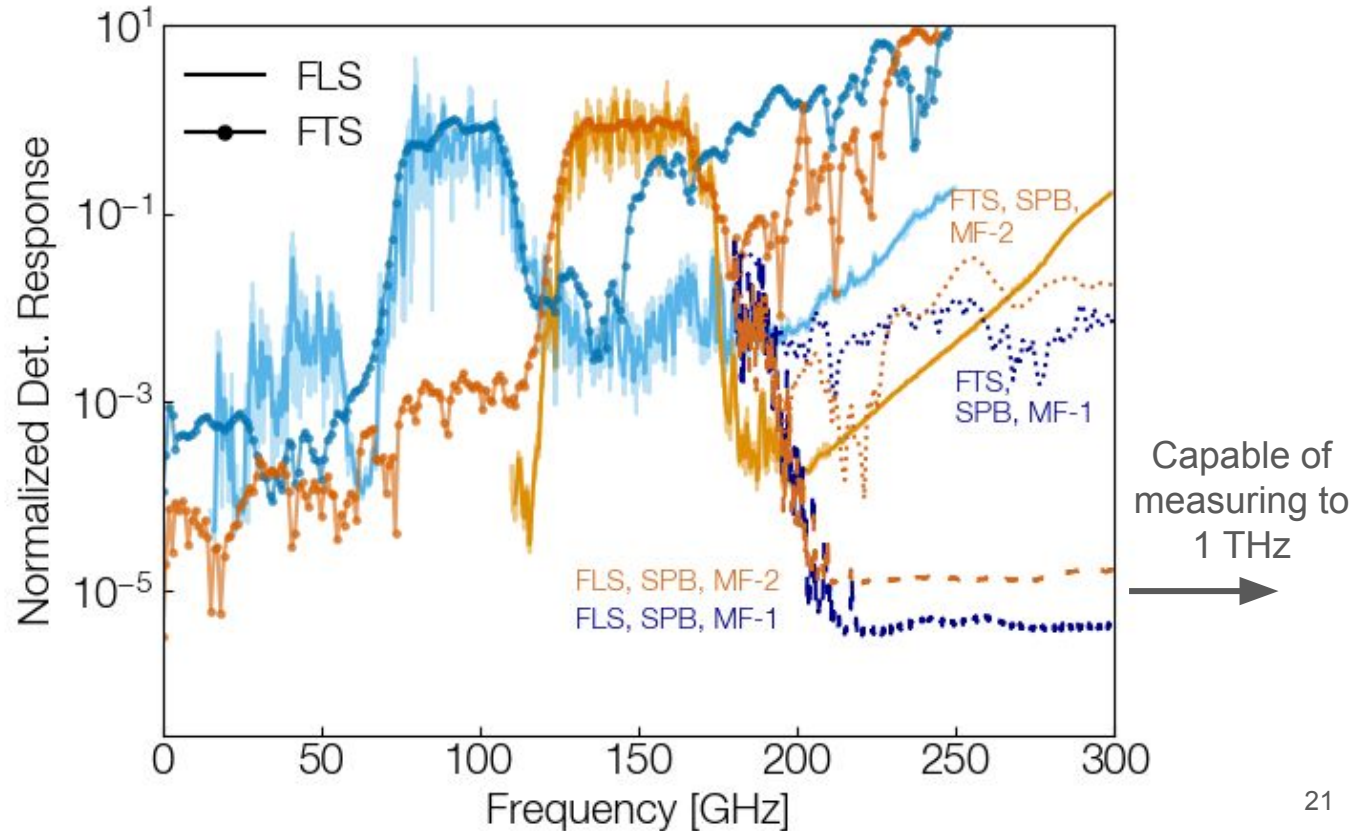
Same periodicity as source power

Divide by power calibration



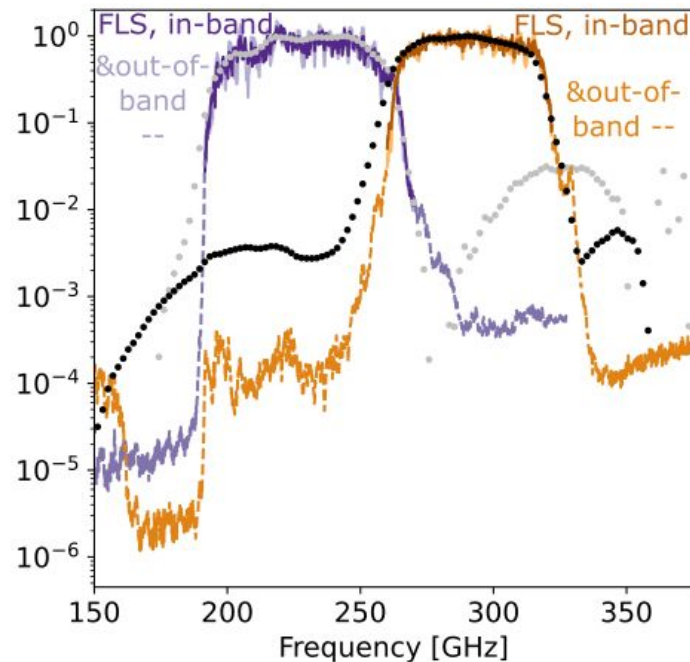
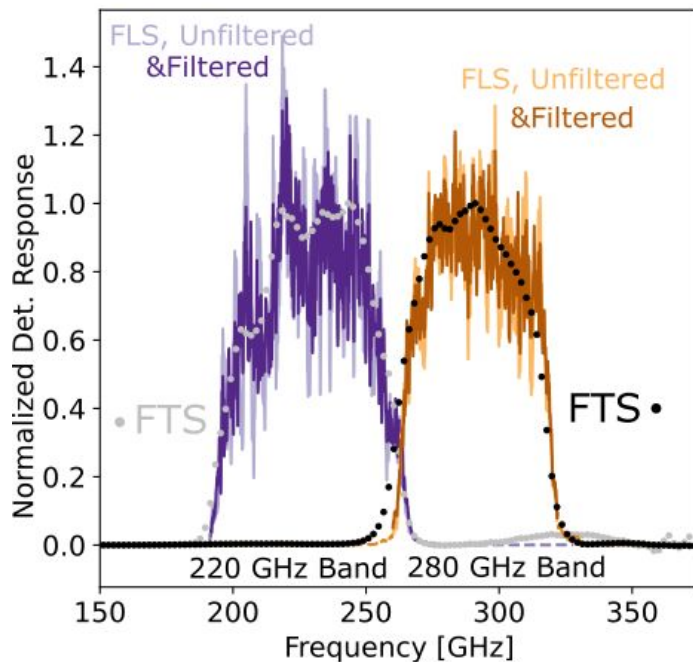
Mid-Frequency LAT Tester Passbands - 90/150 GHz

- FLS and FTS show agreement.
- FLS is more sensitive to band edges and out of band leakage.
- FTS bands are averaged over detector array; FLS bands are individual detectors.



Ultra-High Frequency LAT Tester Passbands - 220/280GHz

- FLS and FTS in agreement wrt band shapes.
- FLS bands smoothed over handful of detectors \rightarrow less noisy band.
- FTS bands are detector array averages.



Summary

- Driving down uncertainties on bandpass measurements is an important and challenging goal.
- The FLS and FTS passbands show agreement and are complementary techniques since they have different systematics.
- Being able to vary the FLS' attenuation allows us to more robustly measure out-of-band and band edge features.
- The FLS offers a promising avenue for bandpass calibration for future experiments.

Thank You!

